Introduction to exoplanetology **Practical work** 2021-2022

# TESS space telescope

Michaël Gillon: michael.gillon@uliege.be Fran Pozuelos: fjpozuelos@uliege.be Mathilde Timmermans: mathilde.timmermans@uliege.be

### TESS (Transiting Exoplanet Survey Satellite)

- Photometric all-sky survey
- NASA mission led by MIT
- Primary mission : search for transiting exoplanets
  - → Earth and super-Earth sized planets



#### Exoplanet in transit – indirect detection method



Drop of brightness when the planet passes in front of the star. Access to planetary and orbital parameters given stellar parameters.

# **TESS** orbit

Credits: NASA

#### Launched in 2018

- Elliptical orbit
- 2:1 resonance with the moon
- P = 13.7 days
- Data transfer at perigee during 16 hours.



### Science instrument

Credits: Ricker et al. (2015)

- 4 cameras arranged in a 4x1 array
- f/1.4 lenses
- CCD detectors
- Red/optical band: 600-1000 nm
- 2048 x 2048 pixels

Optimized for main sequence dwarf stars F5-M5



# Observing strategy

Credits: NASA



### Observing strategy

- Tess Input Catalogue (TIC): more than 200.000 stars.
- Exposure time = 2s.
- Short cadence  $\rightarrow$  2 min.
- Full Frame Images  $\rightarrow$  30 min.
- Two years to scan the whole sky → close to the ecliptic poles: nearly constant coverage.
- 26 sectors of 24° x 96° field of view.



### Observing strategy

- Transit signal detected by the automatic pipeline → TESS Object of Interest (TOI).
- Public data release every 4 months.

All transit-like signals are not due to planets → Need for follow-up observations



## Follow-up observations

TESS Follow-up Observing Program (TFOP) Working Group (WG)

 $\rightarrow$  SG1 : Seeing-limited photometry

- $\rightarrow$  SG2 : Recon Spectroscopy
- → SG3 : High-resolution Imaging with adaptive optics
- → SG4 : Precise Radial Velocity Work
- $\rightarrow$  SG5 : Space-based photometry



#### Possible cases

- Transiting planet on target or around a nearby star.
- Eclipsing binaries : often V shaped + chromaticity.
- $\rightarrow$  On target
- → Background star (blend)
- $\rightarrow$  Bound star (triple system)
- False positives



Credits: Santerne et al. (2013)

#### Current status



Credits: TESS/MIT

#### Important discoveries – some examples

#### TOI-700 d : Earth-sized planet in the habitable zone of M2 dwarf star

Gilbert et al. (2020). The First Habitable Zone Earth-sized Planet from TESS. I: Validation of the TOI-700 System.

#### TOI-1338 b : circumbinary planet

Kostov et al. (2020). TOI-1338: TESS' First Transiting Circumbinary Planet

#### TOI-451 : three planets in a triple system

Newton et al. (2021). TESS Hunt for Young and Maturing Exoplanets (THYME). IV. Three Small Planets Orbiting a 120 Myr Old Star in the Pisces–Eridanus Stream\*

#### TOI-178 : six planets in Laplace resonances

Leleu et al. (2021) Six transiting planets and a chain of Laplace resonances in TOI-178

#### LTT 1445Ab

Winters et al. (2019) Three Red Suns in the Sky: A Transiting, Terrestrial Planet in a Triple M Dwarf System at 6.9 Parsecs

## TRAnsiting Planets and Planetes mals Small Telescopes (60 cm)

	<b>TRAPPIST-North</b>	<b>TRAPPIST-South</b>
Observatory:	Oukaïmeden Observatory, Morocco	La Silla Observatory, Chile
Altitude:	2751 m	2315 m
Latitude:	31°12′33″ N / 31.2061° N	29°15′40′′ S / 29.2546° S
Longitude:	7°52'52'' W / 7.8664° W	70°44'52'' W / 70.7394° W
Camera CCD:	2048 x 2048 pixels	2048 x 2048 pixels
Pixel scale:	0.60"/pixel	0.64''/pixel
Field of view:	20' x 20'	22' x 22'



Credits: E. Jehin

### **TRAPPIST filters**



### TRAPPIST robotic telescopes



Credits: L. Van Laeken

Credits: E. Jehin

### Working with science images

- Images = FITS (Flexible Image Transport System) files.
- Contains all the information about the observation and telescope you need.
- To read them: AstroImageJ or Prose.

Keyword	Value	Comment	
BITPIX	16	number of bits per data pixel	
NAXIS	2	number of data axes	
NAXIS1 NAXIS2	2096	length of data axis 1	
BZERO	32768.0	data range offset	
BSCALE	1.0	scaling factor	
EXPTIME	4.000000000E+001	[ISU 8601] UTC date/time of exposure start [sec] Duration of exposure	
EXPOSURE	4.000000000E+001	[sec] Duration of exposure	
SET-TEMP	-30.00000000000000	CCD temperature setpoint in C	
XPIXSZ	15.00000000000000	Pixel Width in microns (after binning)	
YPIXSZ	15.0000000000000	Pixel Height in microns (after binning)	
XBINNING	1	Binning level along the X-axis	
XORGSUBE	0	Subframe X position in binned pixels	
YORGSUBF	0	Subframe Y position in binned pixels	
READOUTM	'2MHz 1CH'	Readout mode of image	
IMAGETYP	'LIGHT '	Type of image	
SITELAT	-29.256666666666668	geographic latitude of observatory	
SITELONG	-70.73	geographic longitude of observatory	
FOCALLEN	4740.0000000000000	Focal length of telescope in mm	
APTDIA	600.0000000000000	Aperture diameter of telescope in mm	
APTAREA	282743.34669113159	Aperture area of telescope in mm^2	
OBIECT	T0I-2441.01	Target object name	
TELESCOP	'TRAPPIST-S'	Telescope name	
INSTRUME	'TRAPPISTS'	Detector Instrument name	
NOTES	trappist	ubserver name	
FLIPSTAT	1 I		
CSTRETCH	'Medium '	Initial display stretch mode	
CWHITE	3070	Initial display black level in ADUs	
PEDESTAL	0	Correction to add for zero-based ADU	
PIERSIDE	WEST '		
DATE TIME OPC	2021-02-18100:28:17.759	UTC date/time when this file was written	
UT	'00:27:37'	[old format] UTC time of exposure start	
TIMESYS	'UTC '	Default time system	
AIRMASS	1.0863690779321673	Target airmass at mid-exposure	
LONG-OBS	-7.07393944444E+001	[deg +R WGS84] Geodetic latitude	
ALT-OBS	2.3150000000E+003	[metres] Altitude above mean sea level	
OBSERVAT	TRAPPIST	Observatory name	
OBICTRA	6.094036111111105	200 right ascension of target (hours) (2000 right ascension of target (hours)	
DEC	-51.68713560309148	EOD declination of target (degrees)	
OBJCTDEC	-51.680166666666665	J2000 declination of target (degrees)	
EWHM	3.38145400286E+000	picestal Std. color band of image or C=Color [pixels] Mean Full-Width-Half-Max of image star	
ZMAG	2.30243875560E+001	Mag zero point for 1 sec exposure	
EQUINOX	2000.0	Equatorial coordinates are J2000	
CTVPE1	177.9795730411862 'RATAN'	Ideg. 0-360 CCWJ Position angle of plate	
CRVAL1	9.14108348943E+001	X-axis coordinate value	
CRPIX1	1.0480000000E+003	X-axis reference pixel	
CTVPE2	DEC-TAN'	Y-axis coordinate type	
CRVALZ CRPIX2	1.0400000000000000000000000000000000000	V-axis reference pixel	
CD1_1	1.82192318629E-004	Change in RA along X-Axis	
CD1_2	-6.42997098506E-006	Change in RA along Y-Axis	
CD2_1 CD2_2	1.82267159527E-004	Change in DEC along X-Axis	
HIERARCH E	'TRAPPISTS'	Instrument name	
HIERARCH E	(60.A-9901 (A))	ESO program identification	
HIERARCH F	'SCIENCE '	Observation category	
HIERARCH E	'OBJECT '	Observation type	
HIERARCH E	'IMAGE '	Observation technique	
BUNIT	'adu '	Equatorial coordinate system Physical unit of array values	
LST	337.488	LST at start (sec)	
UTC	1657.759	UTC at start (sec)	
MID-085	59263.0191810/03	Modified Julian Date of start of exposure	
DATE-END	'2021-02-18T00:28:17.759'	UTC date/time of end of exposure	
ORIGFILE	'TRAPPS.2021-02-18T00_27_37.759.fits'	Original File Name	
ARCFILE	TRAPPS.2021-02-18T00:27:37.759.fits'	Archive File Name	
DATASUM	324361550	data unit checksum updated 2021-02-18/14/54/21	
JD_SOBS	2459263.519187037	Julian Date at start of exposure	
JD_UTC	2459263.5194185185	Julian Date (UTC) at mid-exposure	
HID_UTC BID_TDB	2459263.5202412284 2459263.5210504164	Bancentric JD (UTC) at mid-exposure	
ALT_OBJ	66.96501695421513	Target altitude at mid-exposure	
AZ_OBJ	168.74198676940208	Target azimuth at mid-exposure	
HA_0BJ	-0.47200075235424954	Target hour angle at mid-exposure	
HISTORY	Previous Filename = TRAPPS.2021-02-18T00:27:37.759 fite	i ai get zenith distance at mid-exposure	
HISTORY	Bias corrected with mbias.fits		
HISTORY	Dark corrected with mdark.fits		
HISTORY	and exposure time scaling factor = 2.66666666666666666		
INDIGKT	principeriodeted with militating		

Lock kowwords Soors

### Image calibration

#### Raw image:

- $\rightarrow$  Very noisy
- $\rightarrow$  Big dust particles or trails

#### **Need for calibration images !**



## Calibration images

#### **1.** Bias subtraction

Removes an offset related to the electronics that is pixel-dependant.

Taken with zero exposure time and shutter closed.

#### 2. Dark subtraction

Removes the dark current produced by thermal excitation of the electrons in the detector.

Taken with several exposure times and shutter closed.

Need to be corrected for bias.

This is then subtracted to the science images.



### Calibration images

#### 3. Flat field division

Corrects for the pixel to pixel sensitivity variation, dust particles, imperfections, and vignetting effect.

Taken at dusk or dawn for uniform light and in different filters.

Must be corrected for bias and darks.

Division of the science images by the corrected median of the flats.



# Science images



## Differential photometry

Goal : remove systematic effects in the target star.

#### Aperture photometry:

Flux measurement inside the aperture by summing the pixel values and subtracting the background contribution (outer ring).



### Differential photometry

- Choose a set of comparison stars which have similar brightness (delta mag) and constant throughout the observation.
- Differential flux :

$$F_{diff} = \frac{F_{T1}}{\sum_{i} F_{c,i}}$$



# Meridian flip

- Encountered with German equatorial mounts.
- Change the side of the mount on which the telescope is to avoid it crashing into the mount.



Credits: https://astronomy.mdodd.com/gem\_movement.html

# Meridian flip

- Encountered for German equatorial mounts.
- Change the side of the mount on which the telescope is to avoid it crashing into the mount.
- Result : Flip of 180° of the dec and ra axes. The positions of the stars on the detector change !
  - $\rightarrow$  Offset in the light curve.





26

# Thanks for your attention ! Do not hesitate to contact us with your questions

Michaël Gillon: michael.gillon@uliege.be Fran Pozuelos: fjpozuelos@uliege.be Mathilde Timmermans: mathilde.timmermans@uliege.be